

# METEOROLOGICAL USES 

OF THE STEREOGRAPHIC HORIZON MAP PROJECTION

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16．Abstract
The stereographic horizon map projection is a generalized form of the polar stereographic projection that permits placement of the map center at any point on the earth．With this projection，it is possible to view translating meteorological systems in one perspective regardless of system location．An algorithm was developed for con－ verting a polar stereographic projection to a stereographic horizon projection．A dis－ tortion analysis showed that linear distortions relative to the map center are less than $10 \%$ within $35^{\circ}$ of great circle arc from the center．Two types of meteorological sys－ tems were examined with meteorological satellite infrared radiation data placed in the new map projection to illustrate its use．The cloud pattern evolution（as seen through an infrared atmospheric window）associated with a rapidly developing extratropical storm was used to depict how the stereographic horizon map projection quantitatively reveals the major centers of cloud development，movement，and dissipation relative to the cyclone center．With the stereographic horizon projection，the classical dynamic features of the jet stream appeared to be present in the appropriate quadrants of a composite of Nimbus 3 infrared（ 6.4 to $6.9 \mu \mathrm{~m}$ ）radiation data surrounding 13 jet stream speed maxima．
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# METEOROLOGICAL USES OF THE STEREOGRAPHIC HORIZON MAP PROJECTION 

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## INTRODUCTION

The analysis of meteorological data is usually performed on some standard map projection, such as the Mercator or polar stereographic projections. Cyclones and other translating weather features are displayed and tracked on the map base. The amount of distortion caused by the geometry of the projection varies as the feature moves, depending on the type of projection and the change in latitude of the moving feature. It is often very advantageous to view translating weather phenomena in a Lagrangian sense, that is, to have the map projection translate with the system. For example, it would be useful to have the pole of a polar stereographic projection located at the center of a tropical cyclone and have the map projection translate with the cyclone. Thus, the map distortions remain the same relative to the center regardless of the location of the cyclone.

One map projection that satisfies this objective is called the stereographic horizon projection (Deetz and Adams, 1944). It is essentially a polar stereographic projection with the property that the pole can be placed at any point on the globe. The movable pole can be placed at a common position relative to a sequence of views of a given system, and hence the weather feature can be analyzed with the same perspective throughout its lifetime. Further, similar weather features can be examined with a common reference point over a lengthy period of record leading to a climatological analysis of these features.

## MAP GEOMETRY

The stereographic horizon projection is defined as the projection of the earth's surface onto a plane tangent to the earth's surface at the map center, the vertex of the projection being the point on the earth diametrically opposite the map center.

Figure 1 shows a point $P$ on the earth that is located on the map by extending the chord $\overline{V P}$ to $P^{\prime}$. Point $C$ on the earth is coincident with $C^{\prime}$ (the center of the map). Any point on earth can be


Figure 1-A side view of the stereographic horizon map projection showing the projection of a point $P$ on a globe onto the projection at point $P^{\prime}$.
specified as the map center, including the North and South Poles, and any type of data for any known location can be placed on the stereographic horizon map base.

A projection extensively used for the display of meteorological information is the polar stereographic projection. The geometry described below transfers the data from a polar stereographic projection (with the projection plane tangent to the earth at the pole) to the stereographic horizon projection.

## Conversion from True Coordinates to Pseudo Coordinates

Figure 2 shows the geometrical relationship between points $C$ and $P$. The true latitude and longitude ( $\phi_{p}, \lambda_{p}$ ) of point $P$ were converted to pseudo latitude and pseudo longitude ( $\beta_{p}, a_{p}$ ), with the map center as the pseudo pole designated by latitude and longitude ( $\phi_{c}, \lambda_{c}$ ).

From Figure 2, $\cos \left(90^{\circ}-\beta_{p}\right)=\mathbf{r} \cdot \mathbf{r}_{0}$, where $\mathbf{r}$ and $\mathbf{r}_{0}$ are vectors extending from the center of the earth to $P\left(\phi_{p}, \lambda_{p}\right)$ and $C\left(\phi_{c}, \lambda_{c}\right)$ respectively. The expressions for $\mathbf{r}$ and $\mathrm{r}_{0}$ are given by

$$
\begin{align*}
\mathbf{r} & =\mathbf{i} \cos \phi_{p} \cos \lambda_{p}+\mathbf{j} \cos \phi_{p} \sin \lambda_{p}+\mathrm{k} \sin \phi_{p},  \tag{1}\\
\mathbf{r}_{0} & =\mathbf{i} \cos \phi_{c} \cos \lambda_{c}+\mathbf{j} \cos \phi_{c} \sin \lambda_{c}+\mathrm{k} \sin \phi_{c}, \tag{2}
\end{align*}
$$

where $\mathbf{i}, \mathbf{j}$, and $\mathbf{k}$ are unit vectors. Therefore,

$$
\begin{equation*}
\cos \left(90^{\circ}-\beta_{p}\right)=\sin \beta_{p}=\cos \phi_{c} \cos \lambda_{c} \cos \phi_{p} \cos \lambda_{p}+\cos \phi_{c} \sin \lambda_{c} \cos \phi_{p} \sin \lambda_{p}+\sin \phi_{c} \sin \phi_{p} . \tag{3}
\end{equation*}
$$

Computations were performed only for $\sin \beta_{p} \geqslant 0$. Thus, no computations were made for $P$ in the opposite hemisphere.

The pseudo longitude of $P, a_{p}$, is computed from the law of sines and law of cosines for a spherical triangle:

$$
\begin{align*}
& \sin a_{p}=\frac{\cos \phi_{p} \sin \left(\lambda_{c}-\lambda_{p}\right)}{\cos \beta_{p}},  \tag{4}\\
& \cos a_{p}=\frac{\sin \phi_{p}-\sin \phi_{c} \sin \beta_{p}}{\cos \phi_{c} \cos \beta_{p}} . \tag{5}
\end{align*}
$$

Computation of both $\sin a_{p}$ and $\cos a_{p}$ is required in order to locate $a_{p}$ in the proper quadrant.


C = CENTER OF MAP, THE PSEUDO POLE
$N=$ TRUE POLE
$0=$ CENTER OF EARTH
P = POINT TO BE LOCATED
$\phi_{p}, \lambda_{p}=$ TRUE LATITUDE, LONGITUDE OF $P$
$\Phi_{c}, \lambda_{c}=$ TRUE LATITUDE, LONGITUDE OF $C$
$\boldsymbol{\beta}_{\mathrm{p}}, \boldsymbol{a}_{\mathrm{p}}=$ PSEUDO LATITUDE, LONGITUDE OF P
$\Varangle \mathrm{COP}=90^{\circ}-\beta$
$\Varangle$ NCP $=a$
Figure 2-Geometrical relationships between the true pole, $N$, the pseudo pole of the stereographic horizon projection, $C$, and an arbitrary point, $P$, on the surface of a globe.

## Grid Print Map Coordinates

A convenient method of displaying meteorological information is to use a computer-generated grid print map. These maps are created by placing the data at equally spaced increments on the map. Thus, in the case of the stereographic horizon projection relative to the center of the map, the $x$ and $y$ coordinates of each data location must be known. With the knowledge of $\beta_{p}$, the map scale factor, and the diameter of the earth, we can now compute the distance $d$ on the stereographic horizon map between points $P^{\prime}$ and $C$ (Figure 3). The relationships for $D$ and $d$ are given by

$$
\begin{align*}
& D=\frac{\text { diameter of the earth }}{\text { map scale factor }}=\text { scaled diameter of the earth }, \\
& d=D \tan \left[\frac{\left(90^{\circ}-\beta_{p}\right)}{2}\right]=D \tan \gamma . \tag{6}
\end{align*}
$$

In the computation of the grid print map coordinates $x$ and $y$ from the center of the map, three different cases are considered:
(1) Map centered at North Pole.
(2) Map centered at South Pole.
(3) Map centered at any other point (stereographic horizon map).

Prior to the calculation of $x$ and $y$ in all three cases, the data within the stereographic horizon map can be rotated about the center by specifying an azimuth, $z$, to be added to the west pseudo longitude, $a_{p}$. Figure 4 diagrams the three cases with a rotational azimuth added. All longitudes in the following discussion, whether pseudo or true, are west longitudes.

## Case 1. North Pole Map

In both polar cases the pseudo latitude and pseudo longitude ( $\beta_{p}, a_{p}$ ) are equal to the true values ( $\lambda_{p}, \phi_{p}$ ). Westerly longitudes increase clockwise from the positive $x$ abscissa which is the starting point by historical convention. Therefore, the expressions for the $x$ and $y$ coordinates of $P^{\prime}$ are
and

$$
\begin{equation*}
x=d \cos \left(a_{p}+z\right) \tag{7}
\end{equation*}
$$

$$
\begin{equation*}
y=-d \sin \left(a_{p}+z\right) \tag{8}
\end{equation*}
$$

Case 2. South Pole Map
Westerly longitudes increase counterclockwise from the negative $x$ abscissa. As a result, the $x$ and $y$ coordinates for $P^{\prime}$ are

$$
D=\frac{\text { DIAMETER OF THE EARTH }}{\text { MAP SCALE }}
$$

$d=$ DISTANCE ON MAP BETWEEN CENTER AND POINT $p^{\prime}$
c = PSEUDO POLE (THE MAP CENTER)
$\boldsymbol{B}_{\mathrm{p}}=$ PSEUDO LATITUDE OF POINT P
$P=$ POINT TO BE LOCATED
$\boldsymbol{\gamma}=\frac{90^{\circ}-\boldsymbol{\beta}_{\mathrm{p}}}{2}$

Figure 3-An edgewise view of the stereographic horizon
projection illustrating the distance $d$ between the map
Figure 3-An edgewise view of the stereographic horizon
projection illustrating the distance $d$ between the map projection center, $C^{\prime}$, and the projected point, $P^{\prime}$.

$$
\begin{equation*}
x=-d \cos \left(a_{p}+z\right) \tag{9}
\end{equation*}
$$

and

$$
\begin{equation*}
y=-d \sin \left(a_{p}+z\right) \tag{10}
\end{equation*}
$$

## Case 3. Stereographic Horizon Map

The longitude was arbitrarily selected to increase clockwise from the positive $y$ ordinate. Thus, the $x$ and $y$ coordinates for $P^{\prime}$ are
and

$$
\begin{equation*}
x=d \sin \left(a_{p}+z\right) \tag{11}
\end{equation*}
$$ same meridian at latitudes $\phi_{2}$ and $\phi_{1}$ and are projected onto the map at $A^{\prime}$ and $B^{\prime}$. The length of the arc $\overline{A B}$ on the globe is to be compared with the line $\overline{A^{\prime} B^{\prime}}$ on the map. Since the length of the entire meridian equals the circumference of the globe, $2 \pi R$, the length of the arc $\overline{A B}$ (hereafter called $M_{g}$ ) would be

$$
\begin{equation*}
M_{g}=2 \pi R\left(\frac{\phi_{1}-\phi_{2}}{360^{\circ}}\right) \tag{13}
\end{equation*}
$$

The distance $\overline{A^{\prime} B^{\prime}}$ on the map is obtained by subtracting $\overline{B^{\prime} C}$ from $\overline{A^{\prime} C}$. It can be seen that $\Varangle A O C=90^{\circ}-\phi_{2}$, and from plane geometry it follows that $\Varangle A D C=1 / 2 \Varangle A O C$. Therefore,

$$
\begin{equation*}
\overline{A^{\prime} C}=2 R \tan \left(\frac{90^{\circ}-\phi_{2}}{2}\right)=2 R \tan \left(45^{\circ}-1 / 2 \phi_{2}\right) \tag{14}
\end{equation*}
$$

and, similarly,

$$
\begin{equation*}
\overline{B^{\prime} C}=2 R \tan \left(45^{\circ}-1 / 2 \phi_{1}\right) \tag{15}
\end{equation*}
$$

Thus, $\overline{A^{\prime} B^{\prime}}$ on the map (hereafter called $M_{m}$ ) would be

$$
\begin{equation*}
M_{m}=\overline{A^{\prime} C}-\overline{B^{\prime} C}=2 R\left[\tan \left(45^{\circ}-1 / 2 \phi_{2}\right)-\tan \left(45^{\circ}-1 / 2 \phi_{1}\right)\right] . \tag{16}
\end{equation*}
$$



Figure 4-The geometry of the calculation of $x$ and $y$ grid print map coordinates corresponding to an arbitrary point $P^{\prime}$ on the three possible types of stereographic projections with a rotational azimuth, $z$, added.

$R=$ RADIUS OF THE EARTH
Figure 5-Distortion of an arc length, $M_{\boldsymbol{g}}$, between points $A$ and $B$ along a meridian on a globe of unit radius when it is projected onto the stereographic horizon map $\left(M_{m}\right)$.

Table 1 gives the lengths of $M_{m}$ and $M_{g}$ for a $5^{\circ}$ arc $\left(\phi_{1}-\phi_{2}=5^{\circ}\right)$ centered at different latitudes, the radius of the globe being set equal to 1.0. Notice that $M_{m}$ is exaggerated as the arc is moved away from the pole, whereas $M_{g}$ remains constant. The degree of exaggeration is expressed as $M_{m} / M_{g}$.

The length of a parallel on the globe with its projected length on a stereographic horizon map was compared. Figure 6 depicts the parallel on the globe at latitude $\phi$, having a radius $r_{g}$, and the projection of this parallel onto the map, with a radius $r_{m}$. The length of the parallel $P_{g}$ on the globe is given by

$$
\begin{equation*}
P_{g}=2 \pi r_{g}=2 \pi(R \cos \phi) \tag{17}
\end{equation*}
$$



Figure 6-The calculation of map distortion around a latitude circle as a function of pseudo latitude by the comparison of the total length of the parallel on a globe of unit radius $\left(P_{g}\right)$ with the total length of the parallel on the stereographic horizon map ( $P_{m}$ ).

From (14), the length of the parallel ( $P_{m}$ ) on the map would be

$$
\begin{align*}
P_{m} & =2 \pi r_{m} \\
& =2 \pi\left[2 R \tan \left(45^{\circ}-1 / 2 \phi\right)\right] \\
& =4 \pi R \tan \left(45^{\circ}-1 / 2 \phi\right) . \tag{18}
\end{align*}
$$

Table 2 gives the values of $P_{g}$ and $P_{m}$ for the latitude of the centers of the same $5^{\circ}$ arcs used in the computation of $M_{g}$ and $M_{m}$. The exaggeration $P_{m} / P_{g}$ is approximately equal to the exaggeration $M_{m} / M_{g}$ along the meridian for all latitudes.

A comparison was also made between the area of a zone on the globe and the corresponding area on the stereographic horizon map. Figure 7 shows the projection of a shaded zone onto the map. The area of the shaded zone on the globe is given by

$$
A_{g}=2 \pi R h
$$

where

$$
\begin{equation*}
h=h_{B}-h_{A}=R\left(\sin \phi_{B}-\sin \phi_{A}\right) . \tag{19}
\end{equation*}
$$

Therefore,

$$
\begin{equation*}
A_{g}=2 \pi R^{2}\left(\sin \phi_{B}-\sin \phi_{A}\right), \tag{20}
\end{equation*}
$$



Figure 7-The comparison of the size of an area, $A_{g}$, on a globe of unit radius with the projected area on the stereographic horizon map, $A_{m}$, illustrating the amount of areal distortion on the map.
and the corresponding shaded area on the map is given by

$$
\begin{equation*}
A_{m}=\pi\left(r_{A}^{2}-r_{B}^{2}\right) \tag{21}
\end{equation*}
$$

Table 2-Comparison of the length of a parallel ( $P_{m}$ ) on the stereographic horizon map with the corresponding parallel length $\left(P_{g}\right)$ on a globe of unit radius.

| Latitude <br> of <br> Parallel | Length on <br> the Globe <br> $\left(P_{\boldsymbol{g}}\right)$ | Length on <br> the Map <br> $\left(P_{\boldsymbol{m}}\right)$ | Exaggeration <br> Factor <br> $\left(P_{\boldsymbol{m}} / P_{\boldsymbol{g}}\right)$ |
| :---: | :---: | :---: | :---: |
| 5 | 6.25929 | 11.51497 | 1.83966 |
| 10 | 6.18774 | 10.54445 | 1.70409 |
| 15 | 6.06911 | 9.64253 | 1.58879 |
| 20 | 5.90423 | 8.79908 | 1.49029 |
| 25 | 5.69451 | 8.00567 | 1.40586 |
| 30 | 5.44141 | 7.25521 | 1.33333 |
| 35 | 5.14620 | 6.54165 | 1.27099 |
| 40 | 4.81321 | 5.85980 | 1.21744 |
| 45 | 4.44289 | 5.20517 | 1.17157 |
| 50 | 4.03876 | 4.57379 | 1.13247 |
| 55 | 3.60390 | 3.96217 | 1.09941 |
| 60 | 7.14161 | 3.36716 | 1.07179 |
| 65 | 2.65540 | 2.78590 | 1.04915 |
| 70 | 2.14899 | 2.24579 | 1.03409 |
| 75 | 1.62622 | 1.65440 | 1.01733 |
| 80 | 1.09107 | 1.09942 | 1.00765 |
| 85 | 0.54762 | 0.54865 | 1.00195 |
| 90 | 0.0 | 0.0 | 1.00017 |

Substituting for $r_{A}$ and $r_{B}$ from (14) and (15), we obtain

$$
\begin{equation*}
A_{m}=4 \pi R^{2}\left[\tan ^{2}\left(45^{\circ}-1 / 2 \phi_{A}\right)-\tan ^{2}\left(45^{\circ}-1 / 2 \phi_{B}\right)\right] . \tag{22}
\end{equation*}
$$

Table 3 gives the values of $A_{g}$ and $A_{m}$ along with the exaggeration factor $A_{m} / A_{g}$ for various values of $\phi_{A}$ and $\phi_{B}$ with $\phi_{B}-\phi_{A}=5^{\circ}$.

It has been shown that, due to the similarity in the exaggeration factor along the meridian and parallel for any point on the map, the shapes of small features on the globe are projected onto the map quite accurately, regardless of their distances from the pole (orthomorphic property). The apparent size of a feature is enlarged as its distance from the pole increases. The exaggeration factor is approximately equal to the product of the meridional and zonal exaggeration factors pertaining to the latitude under consideration. Figure 8 depicts the exaggeration factor, $A_{m} / A_{g}$, as a function of latitudinal distance from the pole of the projection.

Table 3-Comparison of area $\left(A_{m}\right)$ on the stereographic horizon map with area $\left(A_{g}\right)$ on a globe of unit radius.

| Latitudes |  | Area on the Globe$\left(A_{g}\right)$ | Area on the Map ( $A_{m}$ ) | Exaggeration Factor $\left(A_{m} / A_{g}\right)$ |
| :---: | :---: | :---: | :---: | :---: |
| $\phi_{A}$ | $\phi_{B}$ |  |  |  |
| 0 | 5 | 0.54762 | 2,01436 | 3.67933 |
| 5 | 10 | . 54345 | 1.70368 | 3.13494 |
| 10 | 15 | . 53515 | 1.44887 | 2.70743 |
| 15 | 20 | . 52277 | 1.23779 | 2.36776 |
| 20 | 25 | . 50641 | 1.06100 | 2.09513 |
| 25 | 30 | . 48620 | 0.91138 | 1.87448 |
| 30 | 35 | . 46230 | 0.78343 | 1.69465 |
| 35 | 40 | . 43487 | 0.67289 | 1.54734 |
| 40 | 45 | . 40413 | 0.57642 | 1.42632 |
| 45 | 50 | . 37032 | 0.49133 | 1.32677 |
| 50 | 55 | . 33369 | 0.41546 | 1.24506 |
| 55 | 60 | . 29451 | 0.34704 | 1.17836 |
| 60 | 65 | . 25310 | 0.28461 | 1.12446 |
| 65 | 70 | . 20976 | 0.22692 | 1.08177 |
| 70 | 75 | . 16483 | 0.17290 | 1.04895 |
| 75 | 80 | . 11864 | 0.12162 | 1.02512 |
| 80 | 85 | . 07155 | 0.07223 | 1.00956 |
| 85 | 90 | . 02391 | 0.02396 | 1.00189 |
|  |  |  |  |  |

Figure 8-The areal exaggeration factor, $A_{m} / A_{g}$, as a function of pseudo latitude from the center of the stereographic horizon map.

## EXAMPLES OF METEOROLOGICAL USES OF THE PROJECTION

## Tracking of a Cyclone

The great advantage of viewing a weather system in the same perspective as it translates is illustrated by viewing an extratropical cyclone in the stereographic horizon and Mercator map projections on two consecutive days. A western Pacific cyclone center moved from $40^{\circ} \mathrm{N}, 149^{\circ} \mathrm{E}$ at 1400 GMT on November 10, 1966, to $49^{\circ} \mathrm{N}$, $157^{\circ} \mathrm{E}$ twenty-four hours later. Figures 9 and 10 show the National Meteorological Center (NMC) surface charts for 1200 GMT on November 10 and 11,1966 , respectively. The cyclone deepened and intensified rapidly during the 24 -hour interval, the central pressure falling from 1006 mb at the beginning of the period to 987 mb at the end.

In this example, the data to be shown in the stereographic horizon and Mercator map projections are the emitted terrestrial radiation sensed in the 3.5 to $4.1 \mu \mathrm{~m}$ atmospheric "window" by the high resolution infrared radiometer (HRIR) on board the Nimbus 2 meteorological satellite. Since in this wavelength region thermal emission and reflected solar radiation contribute about equally to the observed radiance during the daytime, reasonable meteorological interpretations of terrestrial radiation patterns can be made only at night. Under clear sky conditions, the radiances (which are converted to equivalent blackbody temperatures) are used to determine surface temperatures. Cloud heights can be inferred from the equivalent blackbody temperatures when the clouds are opaque and completely fill the field of view of the radiometer. The interpretation of equivalent blackbody temperatures below 240 K is difficult due to substantial noise in the instrument observations and the method of data processing. A complete error analysis of the Nimbus 2 HRIR data is given by Williamson (1969).


Figure 9-NMC surface chart for 1200 GMT, November 10, 1966.


Figure 10-NMC surface chart for 1200 GMT, November 11, 1966.

Detailed information concerning the Nimbus 2 satellite and the HRIR experiment is available in the "Nimbus II Users' Guide."*

One form of displaying the rectified infrared data is to use the grid print map for various map projections and scales. The algorithms described in the previous section allow the data to be assembled on the stereographic horizon projection in grid print form. More description of the grid print map display can be found in the "Nimbus II Users' Guide" and in other publications, such as Warnecke et al. (1968).

The Mercator and stereographic horizon projections, Figures 11 and 12, respectively, show the radiation pattern composite (orbits 2386 and 2387 ) associated with the extratropical storm for the first day. The frontal positions, adjusted for the time of the satellite pass, were taken from analyses prepared by NMC. Low equivalent blackbody temperatures, clearly depicted near the center of each map, indicate the main cloud mass associated with the cyclone. The shape of the solid cloud deck suggests a storm in the early stage of rapid development (Brodrick, 1969). An indication of dry air intruding into the cloud mass is shown by the notch in the isolines of equivalent blackbody temperature just west of the position of the low center on the surface chart. The position of the stationary front between $25^{\circ}$ and $30^{\circ} \mathrm{N}$ is corroborated by the relatively low equivalent blackbody temperature (cloud) band stretching east-northeast to west-southwest at a substantial distance to the southwest of the cyclone center.

The map scale for each projection is $1: 5,000,000$. Whereas the scale factor is true at the pole (center of the map) for the stereographic horizon projection, it is true at $22.5^{\circ}$ north or south of the equator for the Mercator projections. The cyclone appears smaller in the stereographic horizon projection because of the properties of the Mercator projection with the cyclone center at $40^{\circ} \mathrm{N}$. If the cyclone center were at $22.5^{\circ} \mathrm{N}$, then the system would look similar in both projections. Not only is the apparent map factor different for the two views of the system, but the change in map distortion across the cyclone's cloud mass is greater for the Mercator projection. This results in the impression that a large portion of the cloud deck is north of the cyclone center. A glance at the stereographic horizon map shows that the proportion of the cloud mass north of the center to the total cloud mass is less than that suggested by the Mercator map.

Figures 13 and 14 view the cyclone in the two projections 24 hours later, after the storm has moved some $9^{\circ}$ of latitude farther north. The cloud mass associated with the storm has expanded considerably, especially to the southeast of the storm center, and has assumed the shape of an inverted fishhook. This shape is characteristic of a cyclone that is nearing maximum intensity (Brodrick, 1969).

Since the storm has moved poleward, the differences between the Mercator and stereographic horizon projections are more striking than on the first day. With increasing latitude, the cloud shield size becomes proportionally greater on the Mercator map base. The cloud mass north of the storm center now has become greatly distorted relative to the cloud mass of the entire system.

Table 4 was prepared to illustrate quantitatively the scale difference between the two projections and the distortion change with increasing latitude on a Mercator map. The entries in the table are the number of grid intersections on the grid print maps within the 255 K isotherm, which essentially
*Nimbus Project, NASA Goddard Space Flight Center, Greenbelt, Maryland, 229 pp., 1966.

Table 4-The number of grid intersections (orbits $2399-2400$ ) within the $255-\mathrm{K}$ isotherm bounding the major cloud mass associated with the low pressure center shown in Figures 13 and 14.

| Latitude | Mercator <br> Map Projection <br> $\left(N_{m}\right)$ | Stereographic Horizon <br> Map Projection <br> $\left(N_{s}\right)$ | Comparison Ratio <br> $\left(N_{m} / N_{s}\right)$ |
| :---: | :---: | :---: | :---: |
| $55^{\circ}$ to $60^{\circ} \mathrm{N}$ | 206 | 75 |  |
| $50^{\circ}$ to $55^{\circ} \mathrm{N}$ | 316 | 139 | 2.82 |
| $45^{\circ}$ to $50^{\circ} \mathrm{N}$ | 138 | 68 | 2.27 |
| $40^{\circ}$ to $45^{\circ} \mathrm{N}$ | 90 | 59 | 2.03 |
| $35^{\circ}$ to $40^{\circ} \mathrm{N}$ | 64 | 48 | 1.53 |

bounded the cloud mass associated with the cyclone in Figures 13 and 14. The ratio of the number of grid intersections in the Mercator projection to the number of grid intersections in the stereographic horizon projection within each $5^{\circ}$ latitude band increased with increasing latitude. Between $55^{\circ}$ and $60^{\circ} \mathrm{N}$, the ratio was 2.82 , and between $35^{\circ}$ and $40^{\circ} \mathrm{N}$ the ratio was still substantially greater than unity.

From a meteorological point of view, the most striking chart of this series is an equivalent blackbody temperature difference chart (Figure 15), where the temperatures displayed in Figure 12 (orbits 2386-87) were subtracted from the temperatures depicted in Figure 14 (orbits 2399-2400) in the stereographic horizon projection. It should be emphasized that this difference map was easily calculated with the stereographic horizon projection, whereas it would have been far more difficult to create by other means. Figure 15 essentially reveals the changes in the cloud deck relative to the cyclone center in the $24-\mathrm{hr}$ period. The greatest change occurred to the southeast of the storm center, where the cloud mass developed substantially, as noted by the -60 K differences and the large area covered by differences of more than -50 K . Large positive equivalent blackbody temperature differences ( 20 to 50 K ) indicate the position of the cloud band associated with the subtropical stationary front southwest of the cyclone center on the first day and the subsequent movement of the band from that position relative to the storm center. From the conventional data, it is most likely that the polar cold front further north on the first day merged with the subtropical stationary front, and this combined front swept southward. The smaller negative equivalent blackbody temperature differences south of the axis of the positive equivalent blackbody temperature differences noted above indicate that the cloud band weakened.

Near the low center is another area of large positive equivalent blackbody temperature differences, which are associated with the dry air that encircles the center. The smaller positive differences north-northwest of the center are the result of higher equivalent blackbody temperatures on the second day, a common occurrence as cyclones reach peak intensity. East of this combined area of positive equivalent blackbody temperatures lies a large area of negative equivalent blackbody temperatures, which reflects the expansion of the cloud shield east of the storm center.

In general, the equivalent blackbody temperatures were lower on the second day (Figure 14). As a result, most of the differences shown in Figure 15 are negative. Since the cyclone was located $9^{\circ}$ of


Figure 11-Infrared ( 3.5 to $4.1 \mu \mathrm{~m}$ ) radiation equivalent blackbody temperature ( $K$ ) pattern associated with the extratropical storm centered at $40^{\circ} \mathrm{N}, 149^{\circ} \mathrm{E}$ shown in a Mercator map projection. Stippled areas indicate where the equivalent blackbody temperatures are less than 250 K .
latitude farther north on the second day, the ocean temperature background beneath the clouds was lower than on the first day. Although with HRIR data there is no way to test independently for the presence or absence of clouds, the large area of negative values ( 1 to 5 K ) some $20^{\circ}$ of latitude south of the storm center probably is representative of this effect. This amount of ocean temperature change is consistent with the climatological ocean temperatures for November presented by LaViolette and Seim (1969).


Figure 12-The same equivalent blackbody temperature ( $K$ ) pattern shown in Figure 11 in the stereographic horizon map projection. The concentric circles are pseudo co-latitude distances from the map center. Stippled areas indicate where the equivalent blackbody temperatures are less than 250 K .

## Jet Stream Analysis

The stereographic horizon projection can be used to ascertain if consistent patterns relative to certain features of jet stream flow are discernible in the equivalent blackbody temperature patterns observed by satellite radiometers. As an illustration of this capability, a composite of equivalent


Figure 13 -Infrared ( 3.5 to $4.1 \mu \mathrm{~m}$ ) radiation equivalent blackbody temperature ( $K$ ) pattern associated with the extratropical storm centered at $49^{\circ} \mathrm{N}, 157^{\circ} \mathrm{E}$ shown in the Mercator map projection.


Figure 14 -The same equivalent blackbody temperature ( $K$ ) pattern shown in Figure 13 in the stereographic horizon map projection. The concentric circles are pseudo co-latitude distances from the map center.
blackbody temperatures in the Nimbus 2 medium resolution infrared interferometer (MRIR) 6.4 to $6.9 \mu \mathrm{~m}, 10$ to $11 \mu \mathrm{~m}$, and 5 to $30 \mu \mathrm{~m}$ spectral response regions were assembled on a $1: 10$ million stereographic horizon map for appropriate situations occurring during the period May 20 to July $28,1966$. The composites were formed from 23 orbits and provided data coverage for 13 individual jet stream cases occurring over or near the continental United States. The locations of the jet maxima ranged between $29^{\circ}$ and $47^{\circ} \mathrm{N}$ and $82^{\circ}$ and $115^{\circ} \mathrm{W}$.


Figure 15-Equivalent blackbody temperature ( $K$ ) difference chart where the temperatures displayed in Figure 12 (orbits 2386-87) are subtracted from the temperatures shown in Figure 14 (orbits 2399-2400). The concentric circles are pseudo co-latitude distances from the map center.

In each of the jet stream cases, the stereographic horizon grid was centered at a location corresponding to that of a concomitant wind speed maximum. The location of the jet stream and the associated wind speed maximum were determined from $200-\mathrm{mb}, 300-\mathrm{mb}$, maximum wind speed, and wind shear analyses prepared by NMC. Because the Nimbus orbits do not cross the United States at the standard 0000 and 1200 GMT map times associated with the synoptic analyses just mentioned, it was necessary to locate the wind speed maximum at two successive map times surrounding the MRIR observation time. Once this was accomplished the position of the speed maximum was adjusted to correspond to the time of satellite passage. The difficulty associated with establishing the positions of a wind speed maximum at two successive times from only the NMC analyses resulted in only 13 cases


Figure 16-Composited equivalent blackbody temperatures measured in the 6.4 to $6.9 \mu \mathrm{~m}$ spectral region and associated with 13 jet stream cases where the stereographic horizon map center is located at the wind speed maximum. In each individual jet stream case, a constant azimuth was added to each data point so that all data would be assembled relative to the $x$-axis of the stereographic horizon projection.
being selected from the May 20 to July 28 period. Once the speed maximum was located, a constant azimuth, $z$, was determined for each individual jet stream case that was added to the pseudo longitude $a_{p}$ of each point in the stereographic horizon map. The magnitudes of the azimuth were decided by determining the amount that the data should be rotated in order to bring all data parallel to the direction of flow through the wind speed maximum such that these data were parallel to the $x$-axis of the stereographic horizon projection (see Figure 4). The magnitude of $z$ ranged from $92^{\circ}$ to $-44^{\circ}$. The purpose of this procedure was to form a composite of the observed equivalent blackbody temperatures relative to the jet stream flow passing through the wind speed maximum.

Figure 16 shows the results arising from the compositing procedure in the case of the 6.4 to 6.9 $\mu \mathrm{m}$ MRIR water vapor channel. The response in this channel is related to the amount of moisture in the upper troposphere; i.e., the magnitude of the equivalent blackbody temperatures varies inversely with the amount of moisture. As can be seen in Figure 16, a band of warmer temperatures extends from a position to the rear and left of the speed maximum into the lower right quadrant. There is also a band of lower equivalent blackbody temperatures to the right of the jet, primarily in the lower left quadrant, and another area of lower temperatures in the upper right quadrant. The relative locations of these warmer and colder equivalent blackbody temperature areas, which are most likely associated with convergence and divergence, respectively, are in agreement with results predicted by an advective approximation of the vorticity equation (Reiter, 1963). The pronounced band of colder temperatures to the right of the jet is associated with higher amounts of moisture and clouds of the cirrus type. In general, the results described here are in agreement with studies reported by Johnson and Shen (1967) and Beran et al. (1968) involving radiometric measurements and with studies such as those by Reiter and Whitney (1969) and Whitney et al. (1966) utilizing satellite camera observations.

The reasonable nature of the results described above and the contribution that the horizontal stereographic projection makes in facilitating the consistent assembling of the radiometric satellite data indicate that this technique can be very useful for further studies relating quasi-horizontal fields of emitted radiation to the dynamic characteristics of the jet stream.

## CONCLUDING REMARKS

The horizontal stereographic map projection has considerable value for meteorological analysis. Normalization of the viewing perspective and different orientations of translating meteorological systems can be achieved. The projection permits the compositing of data such that the climatologies of translating systems can be examined more efficiently. Further, statistical analysis can be performed more easily where measurements for weather systems are consistently located in the same perspective. Undoubtedly, other meteorological applications for the use of the stereographic horizon projection will emerge as the Lagrangian property of viewing meteorological systems is exploited.

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